

# The Radiation Explorer in the Far InfraRed prototype: technical overview and measurement performance

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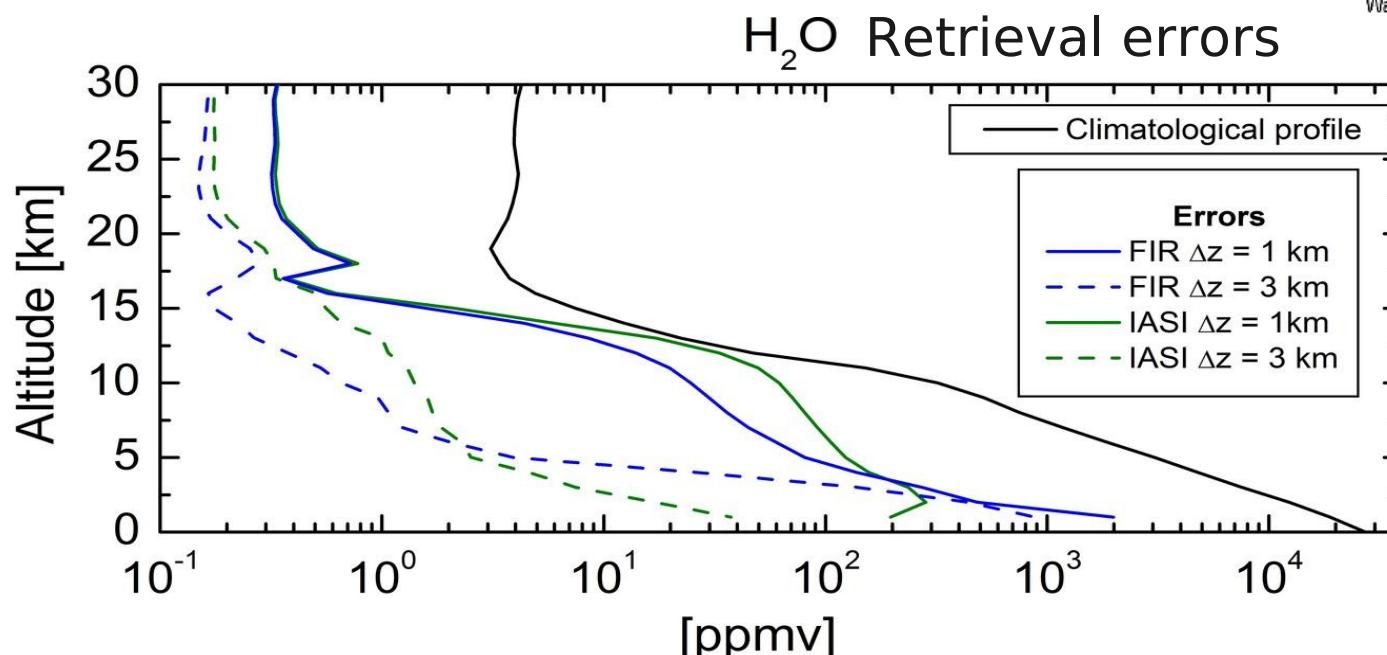
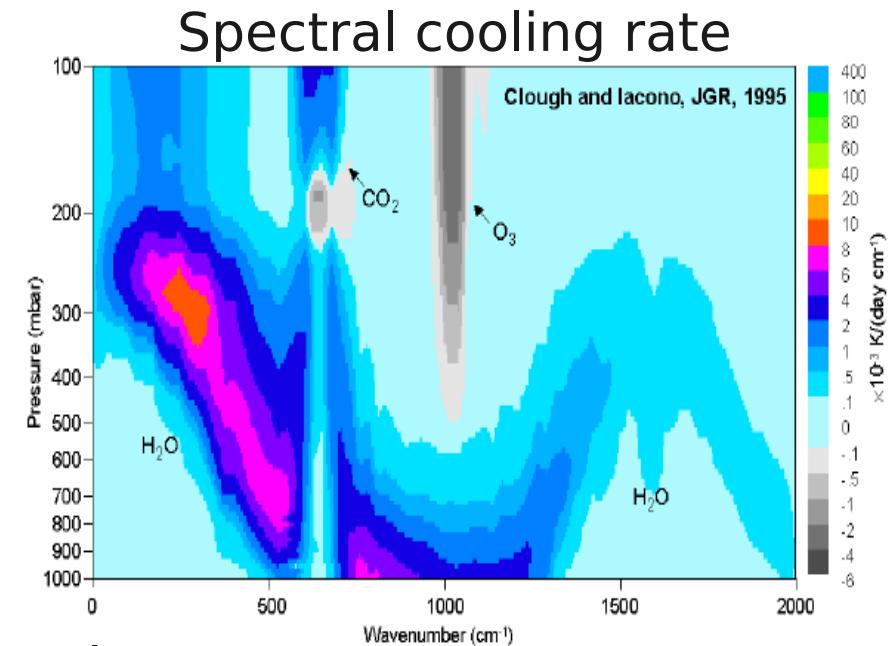
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# Importance of the FIR spectral region

## Water vapour and cirrus

- The FIR region can provide unique information on WV and cirrus clouds radiative effects with sensitivity to microphysics
- The FIR is strongly sensitive to mid-upper level tropospheric humidity that produces a peak in the cooling rate of the atmosphere



FORUM/EFTWVAC  
ESA proposals

# Historical overview

## The REFIR project

### Radiation Explorer in the Far InfraRed

Spectral characterization of the emitted radiance in the 100-1100 cm<sup>-1</sup> range

- REFIR feasibility space mission, EU 1997-2000
- REFIR feasibility for Small Scientific Missions, ASI 1998-2001
- BB pre-development, ESA 2001-2003
  - ADGB-Dip. Fisica, Univ. Bologna, Italy  
PI Rolando Rizzi
  - Selex-Galileo, Campi Bisenzio, Italy
  - Blackett Laboratory, Dep. Physics, ICSTM, London, UK
  - CNRS-SNCMP, Toulouse, France
  - Space Science Dep., CCRLC-RAL, Chilton, UK
  - DIFA, Univ. Basilicata, Potenza, Italy
  - Dip. Fisica, Univ. La Sapienza, Roma, Italy



- REFIR-BB laboratory applications, UniBas, IFAC 2000-2001
- REFIR-PAD ground-based and stratospheric balloon, IFAC 2003-present
  - IFAC-CNR, Firenze, Italy
  - DIFA, Univ. Basilicata, Potenza, Italy
  - IMAA-CNR, Potenza, Italy
  - ADGB-Dip. Fisica, Univ. Bologna, Italy



- Recent proposals for space missions:
  - FORUM
  - EFTWVAC

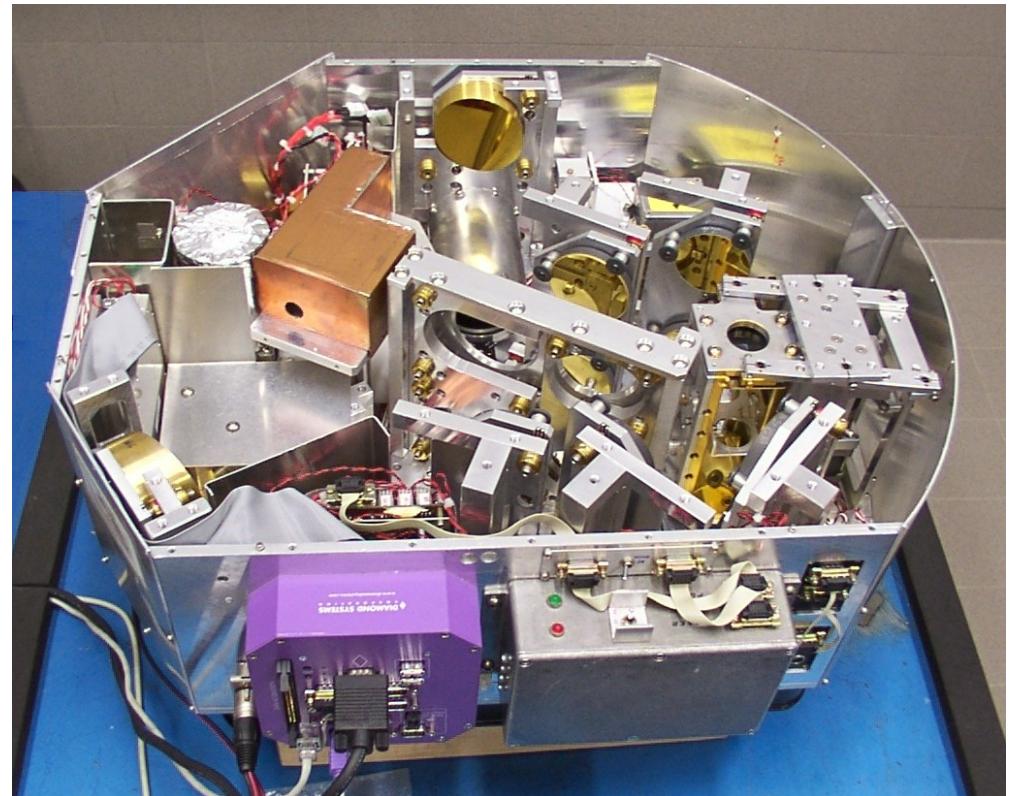
REGIONE  
TOSCANA



# The REFIR-PAD prototype

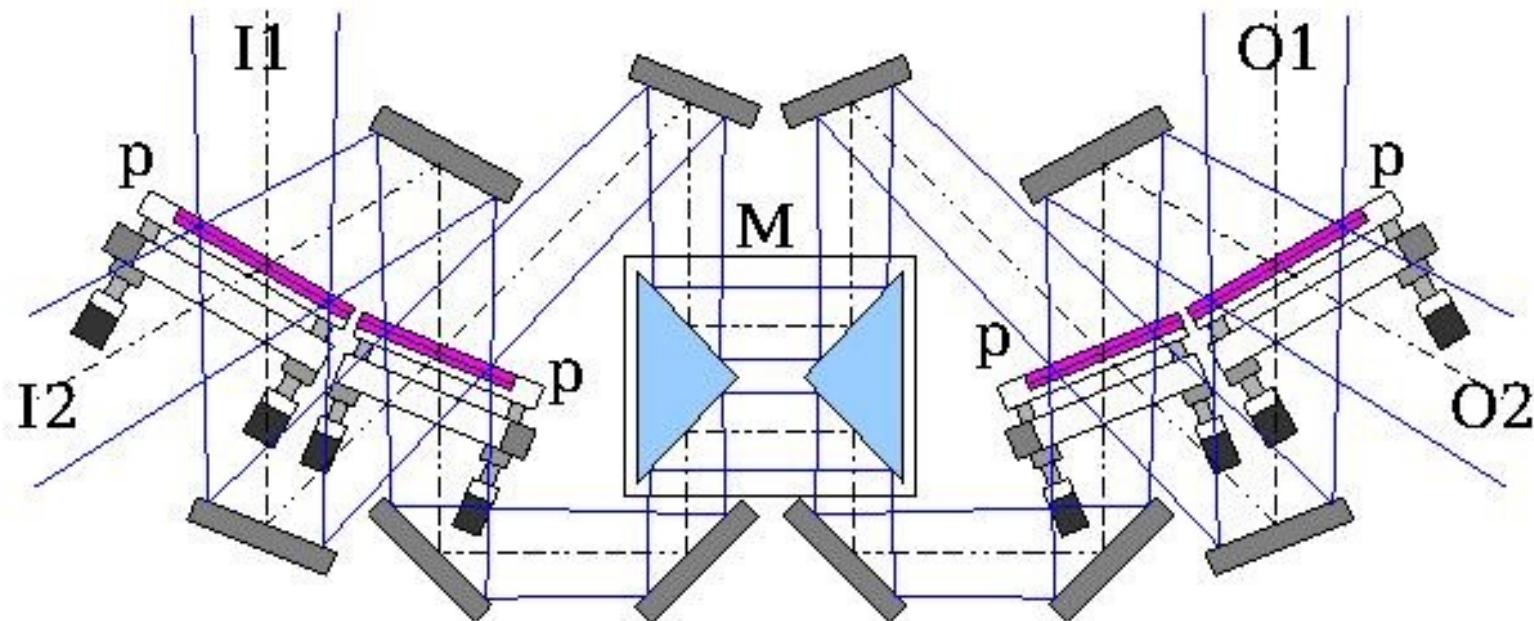
## Radiation Explorer in the Far-InfraRed – Prototype for Applications and Development

- Developed in 2003-2004
- Fourier Transform Spectrometer
- Broadband Ge on PET beam splitters and DLATGS pyroelectric room-temperature detectors
  - Spectral coverage = 100-1400  $\text{cm}^{-1}$ ,
  - Resolution 0.25  $\text{cm}^{-1}$  max. double-sided
  - NESR in the range 0.8-2.5  $\text{mW}/(\text{m}^2 \text{ sr cm}^{-1})$  with 30 s. acquisition time
- Small Payload: 62 cm dia., 55 kg weight, 50 W avg power



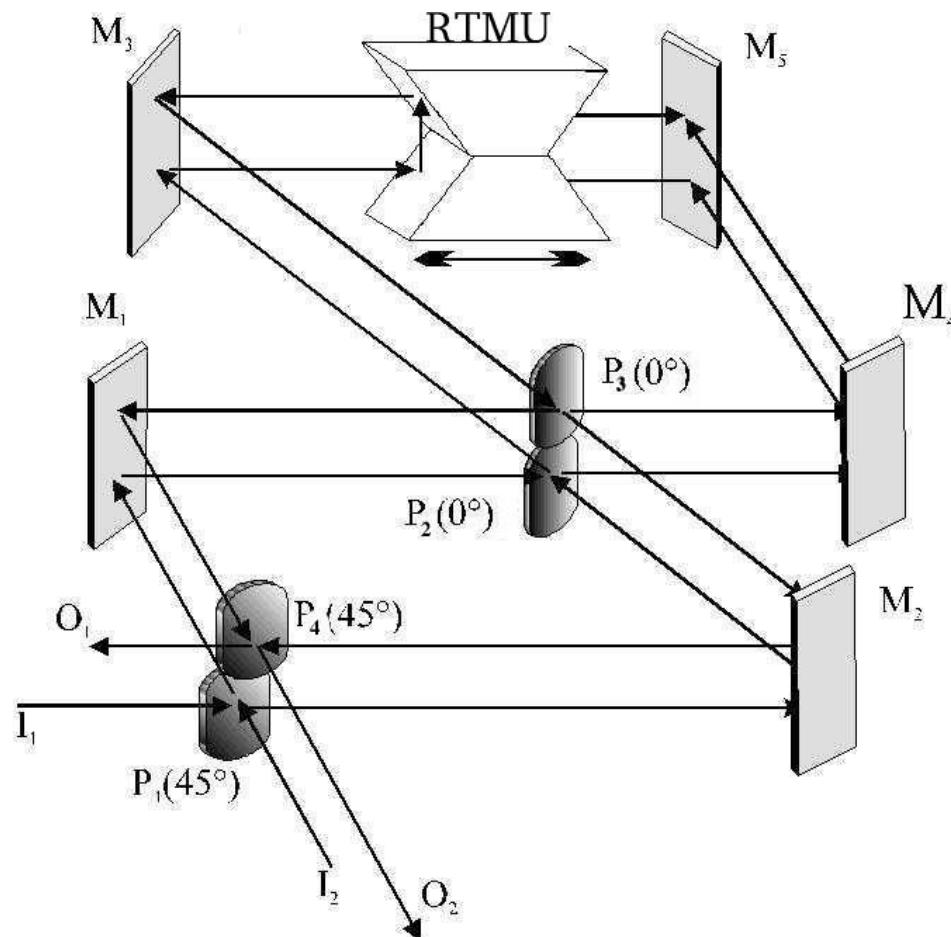
# Optical configuration of interferometer

- 2 input / 2 output ports
- Martin-Puplett (polariser BSs) or Mach-Zehnder (amplitude BSs) interferometer
- Tilt and lateral shift compensation of moving mirrors
- 4x folding OPD



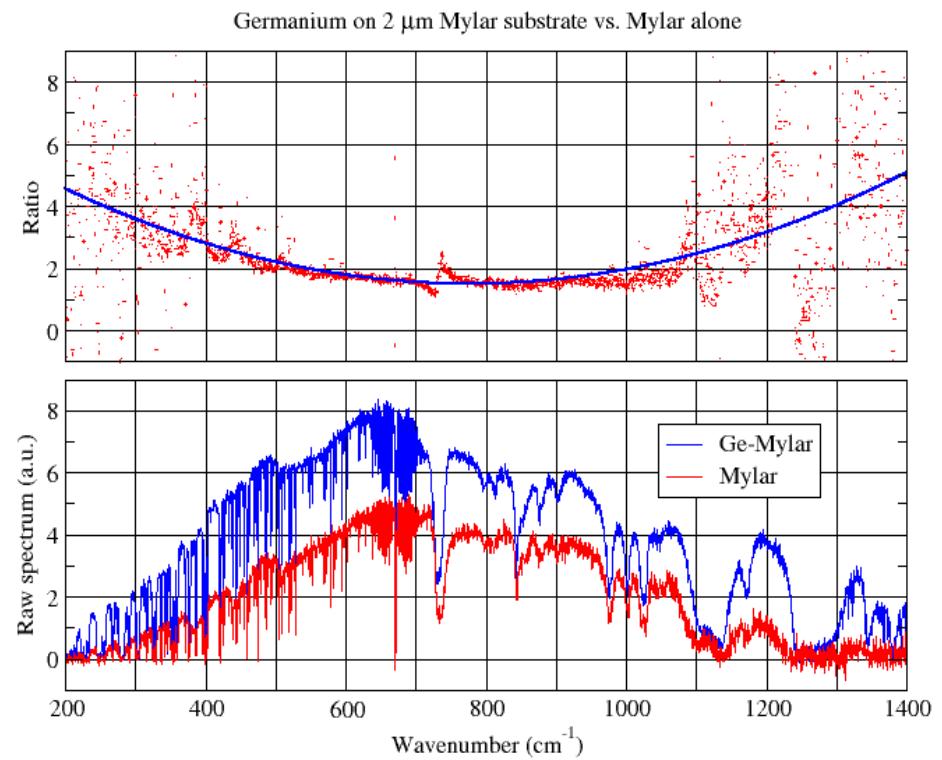
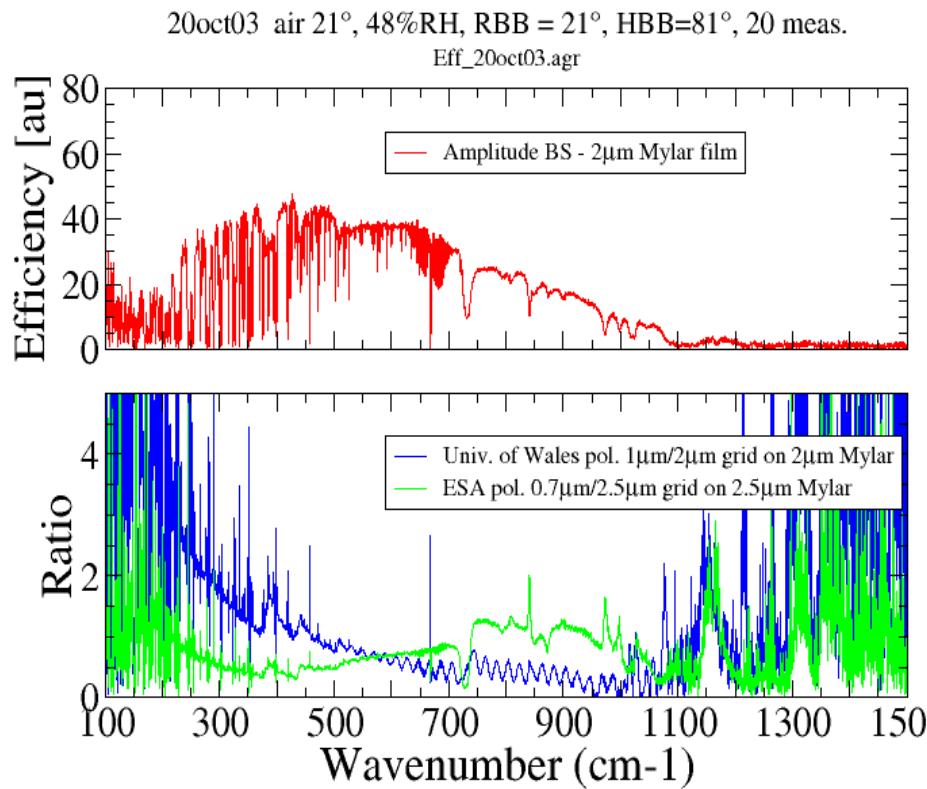
# Two levels interferometer

- Folding around the moving mirror to obtain two levels containing input ports (lower level) and output ports (upper level)
- Horizontal tilt compensation



# Broadband Beam Splitters

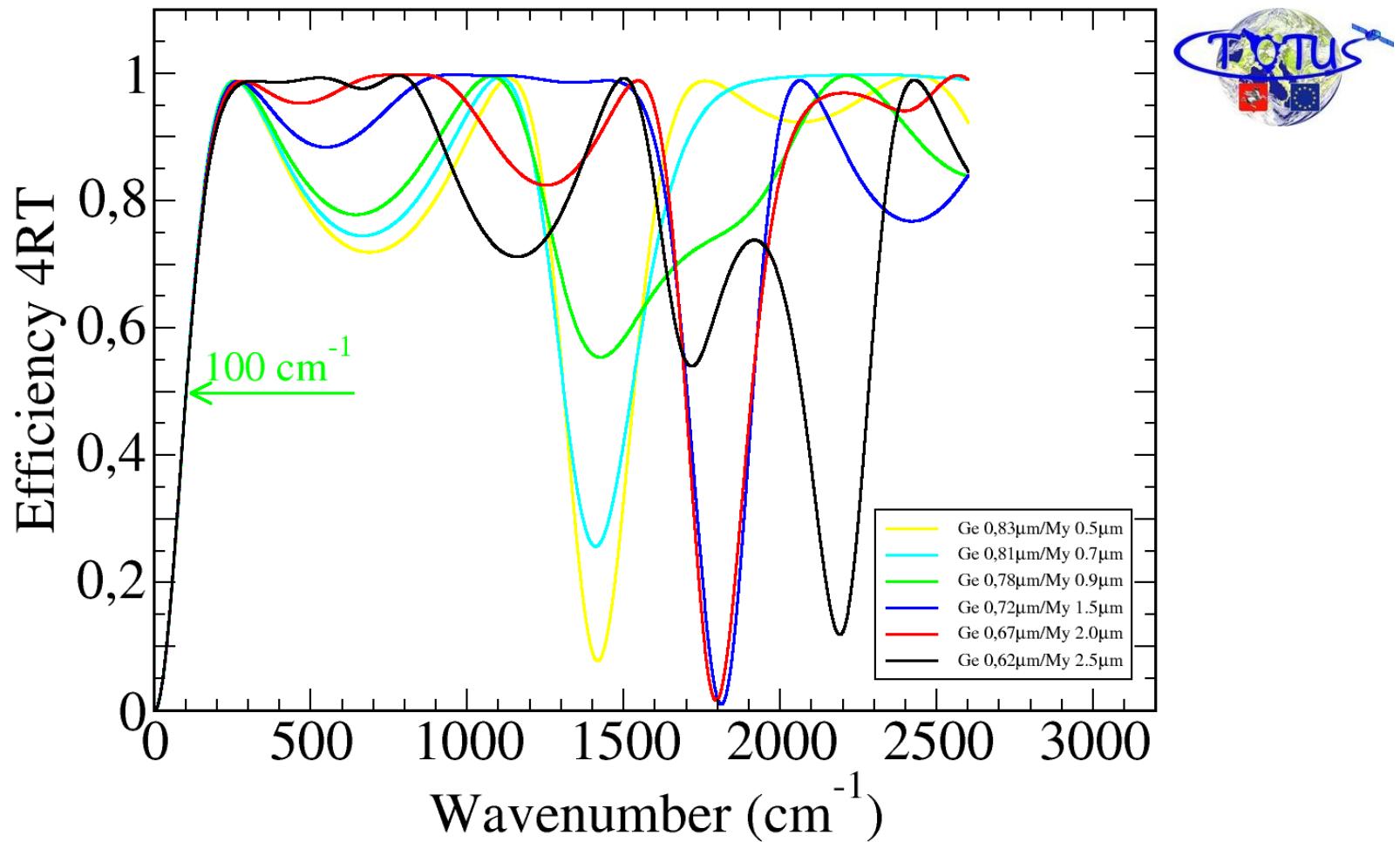
- Photolithographic wires on PET
- Single PET/PP layer
- Ge-PET/PP bilayer



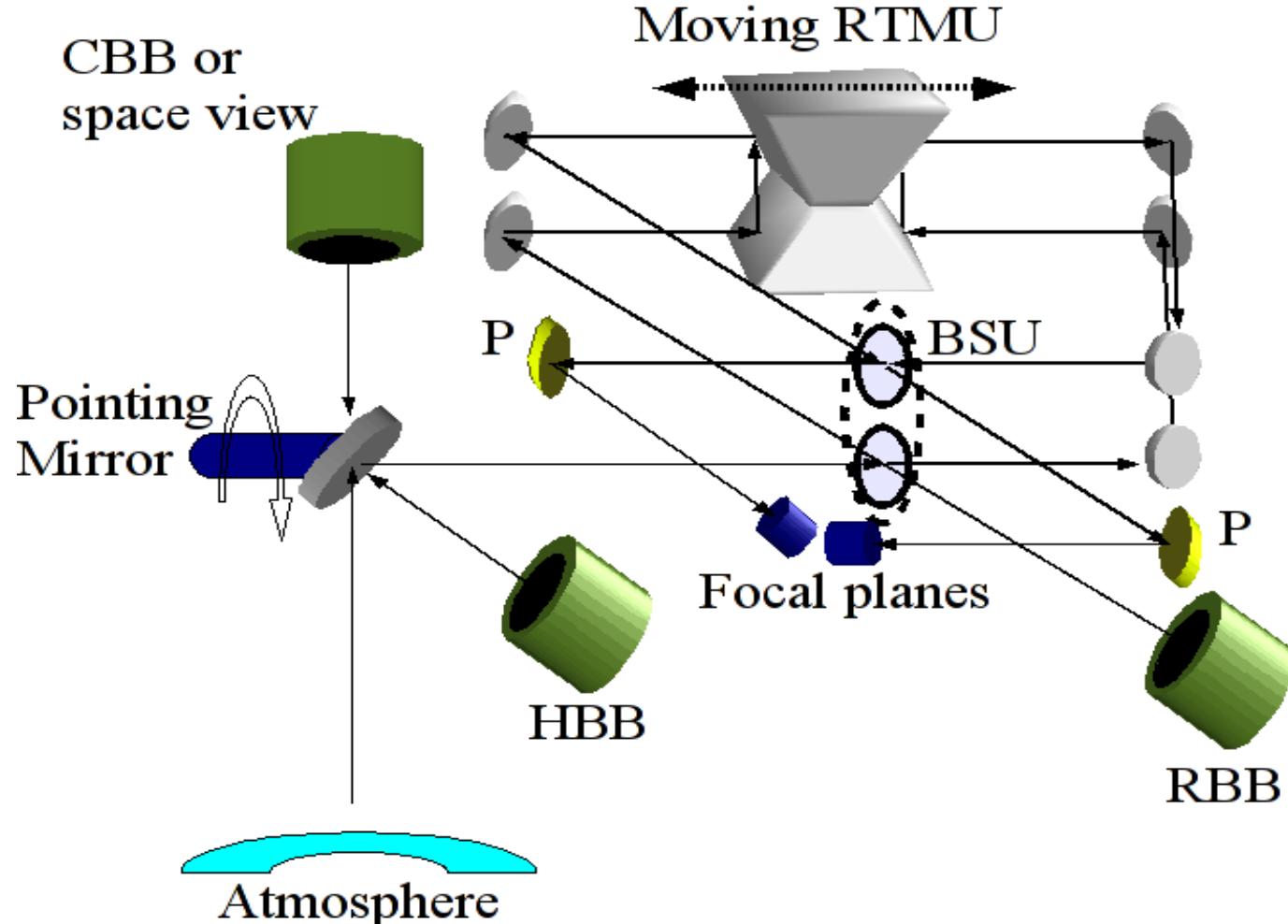
# Wideband beam-splitters design

- Bilayer structures

BS Efficiency estimated for FORUM, Mylar-Ge, 25 deg of incidence  
eff\_mylar-germanium\_OSA-FTS.agr

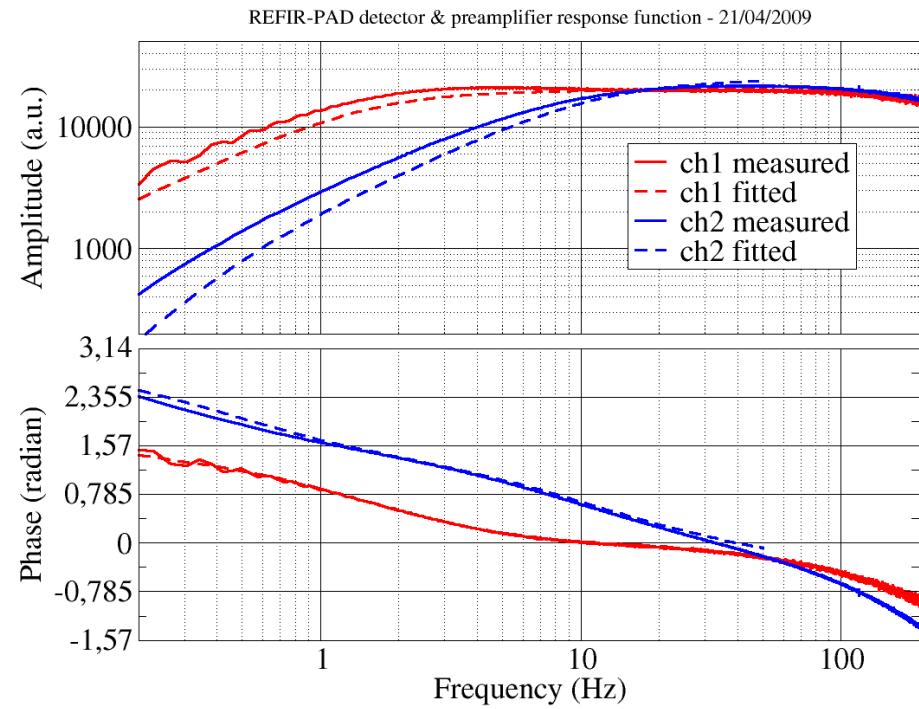
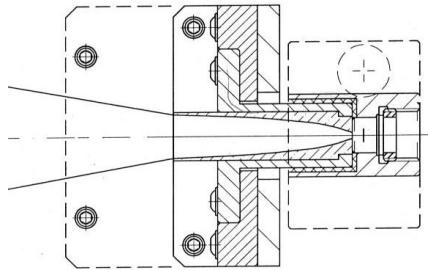


# REFIR-PAD FTS optical layout



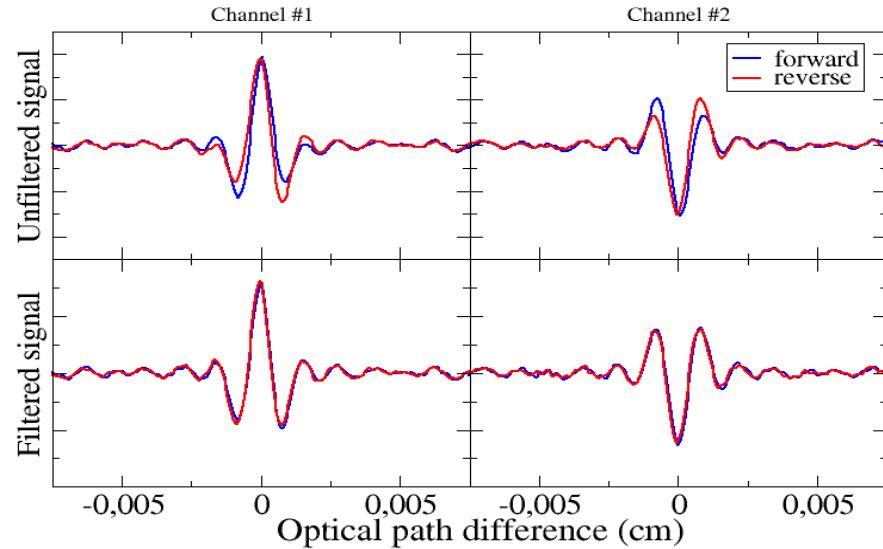
# Interferogram Sampling

DLATGS pyroelectrics  
Winston cones



Equal Time Sampling

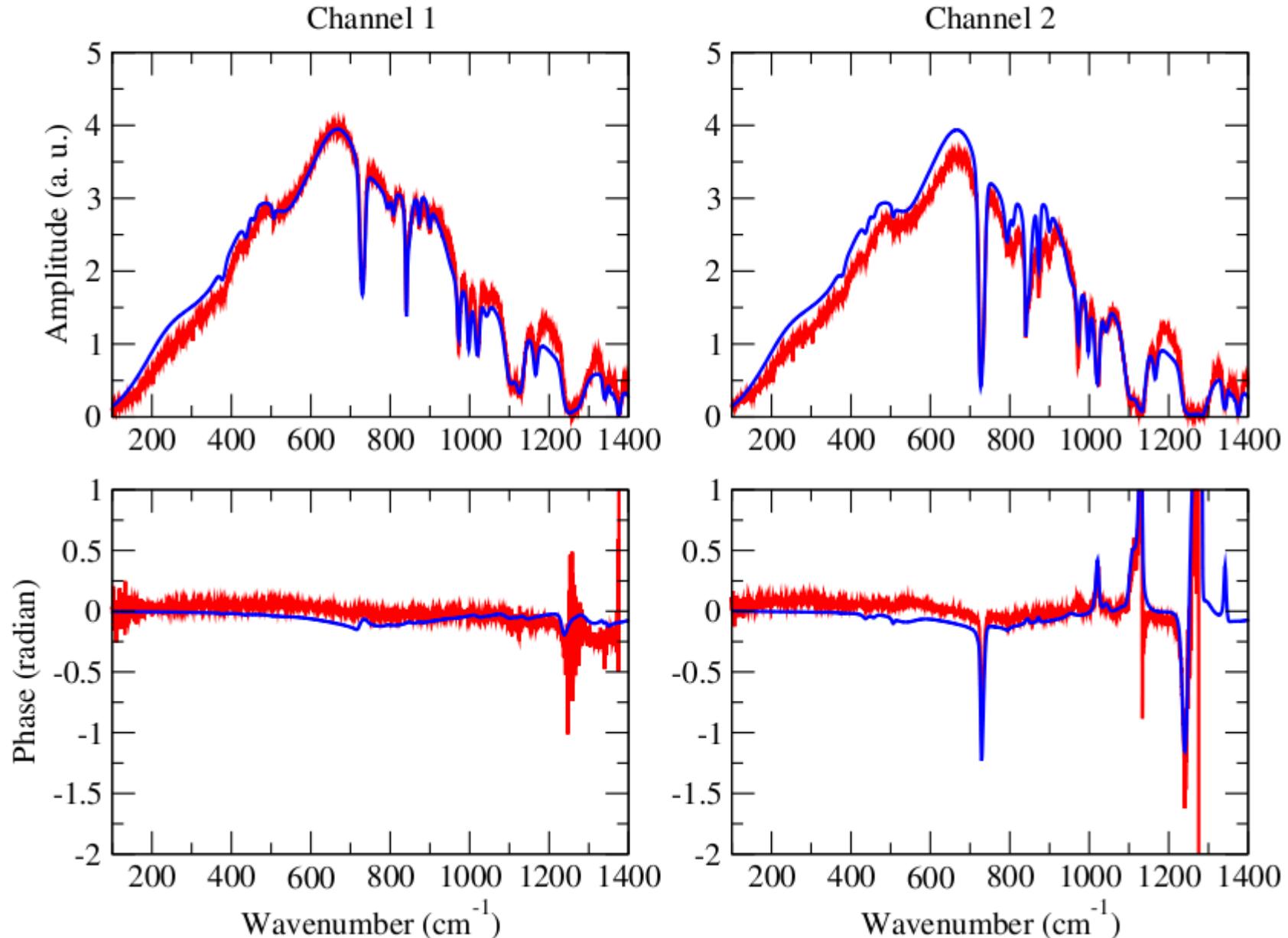
Filtering and Resampling



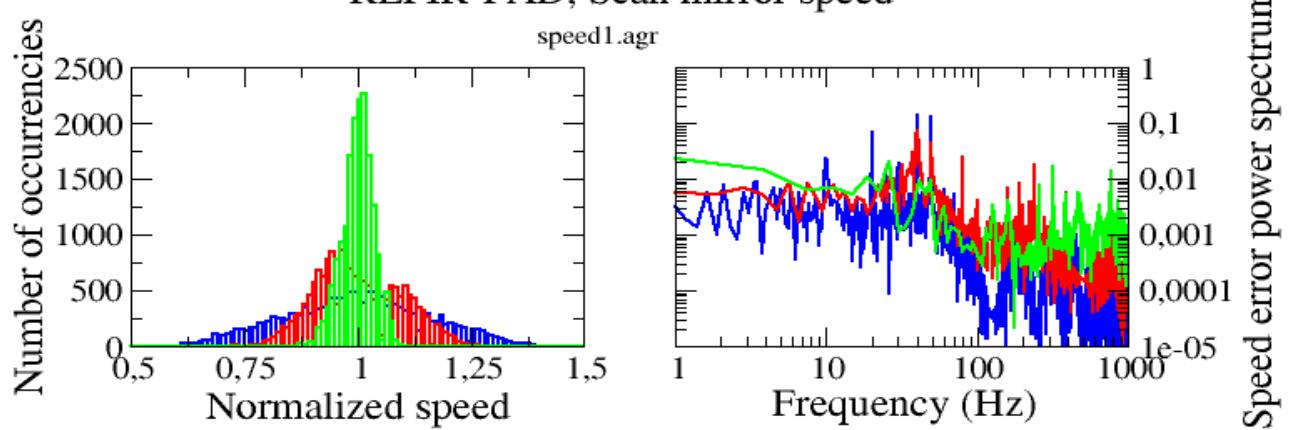
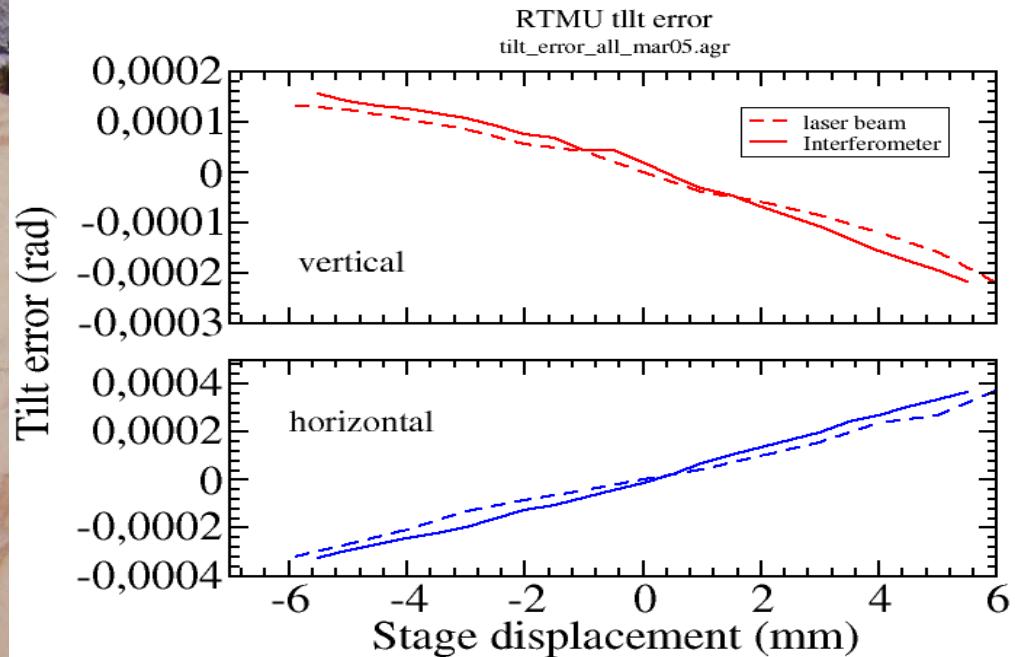
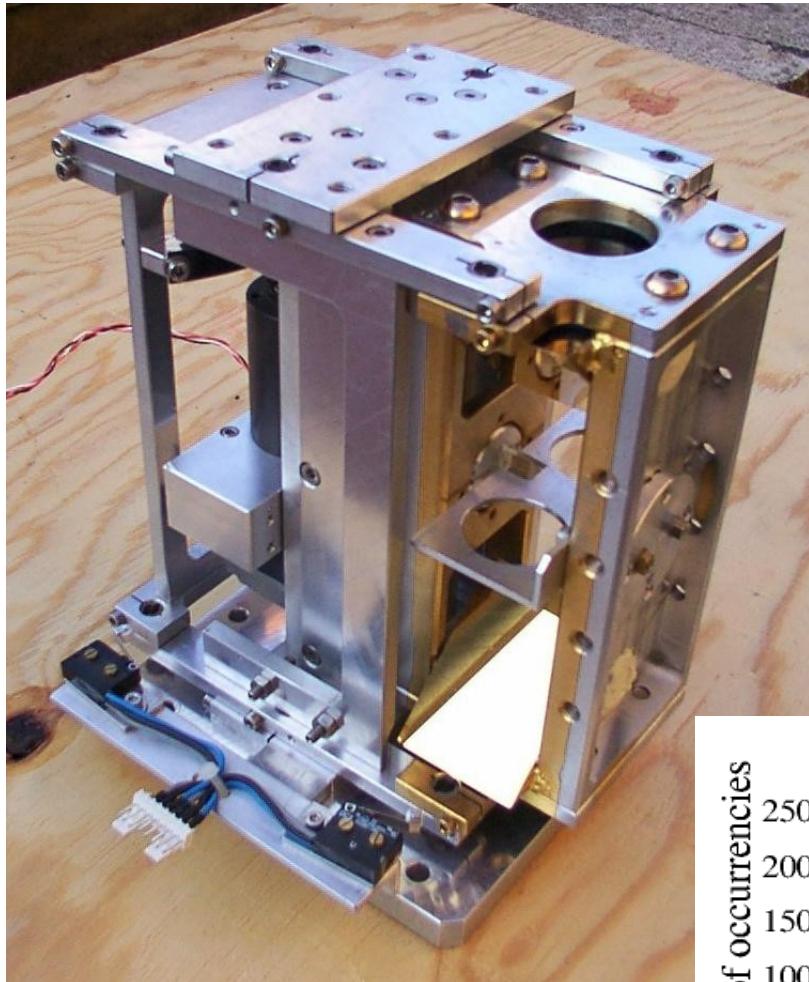
# Uncalibrated Spectra

Hot blackbody source ( $T_M = 347$  K)

$T_R = 302$  K,  $T_{BS} = 302$  K



# Scanning mirror performances



# Radiometric calibration

- Uncalibrated complex spectrum  $S(\sigma)$  given by:

$$S(\sigma) = F1(\sigma)L(\sigma) - F2(\sigma)L_{RBB}(\sigma)$$

- Calibration with 2 views of reference sources: HBB,CBB. With  $L_{RBB}$  is constant we have:

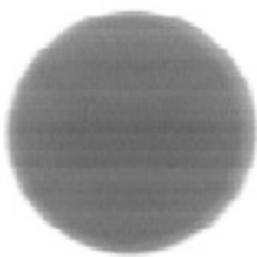
$$L(\sigma) = \mathbb{R} \left\{ \frac{S(\sigma) - S_{CBB}(\sigma)}{S_{HBB}(\sigma) - S_{CBB}(\sigma)} \cdot (L_{HBB}(\sigma) - L_{CBB}(\sigma)) + L_{CBB}(\sigma) \right\}$$

- The used optical layout allows the best way to have access to all the sources and with minimum phase errors

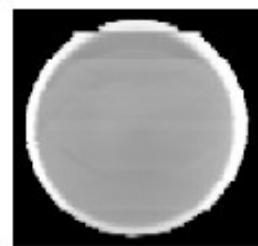
# Black-body calibration sources

- RBB for reference input
- HBB/CBB for calibration
- Cylinder + cone cavity
- Xylan<sup>©</sup> coated surface
- Emissivity estim.  $\sim 0.99$
- PT100 temperature meas.
- $\Delta T < 0.3$  K

Thermal images



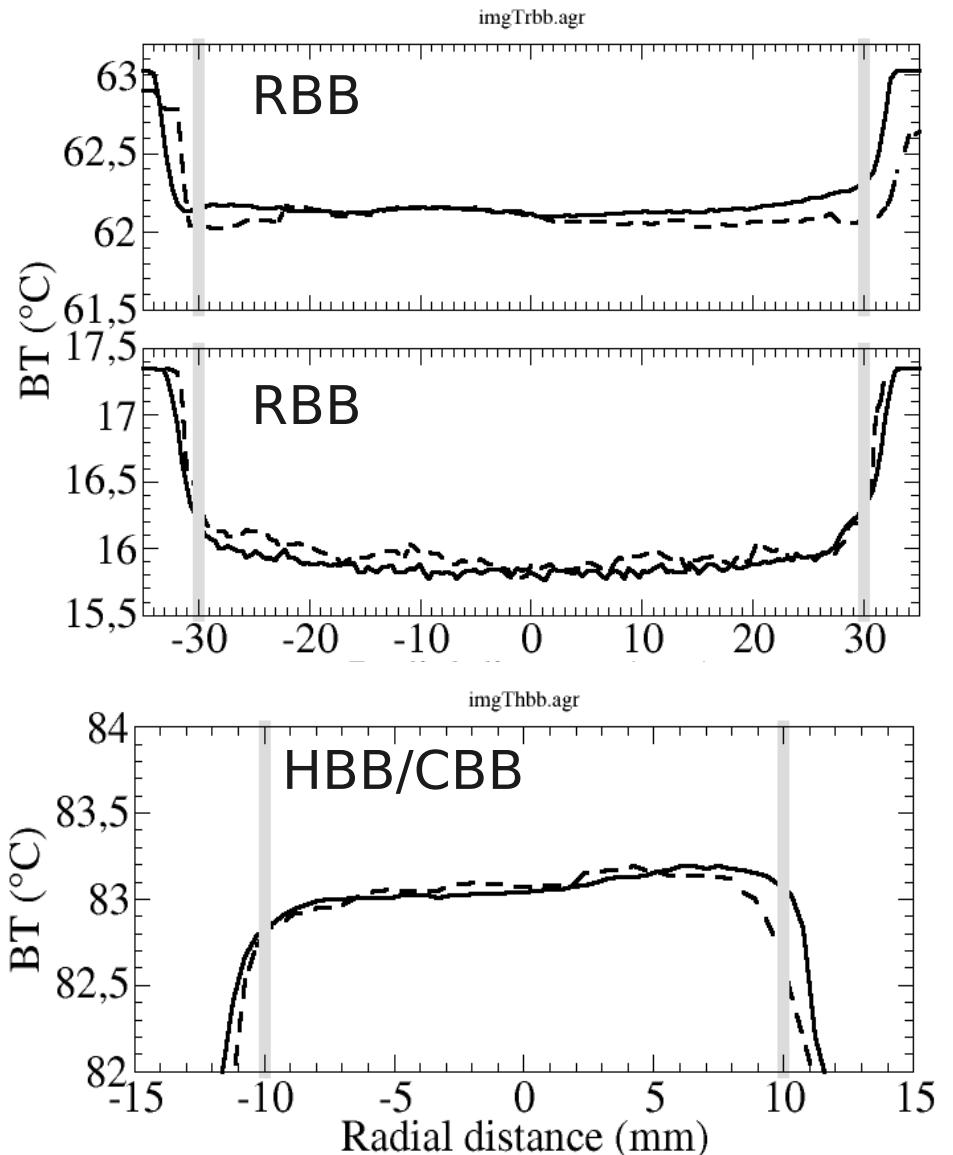
RBB  
cooled



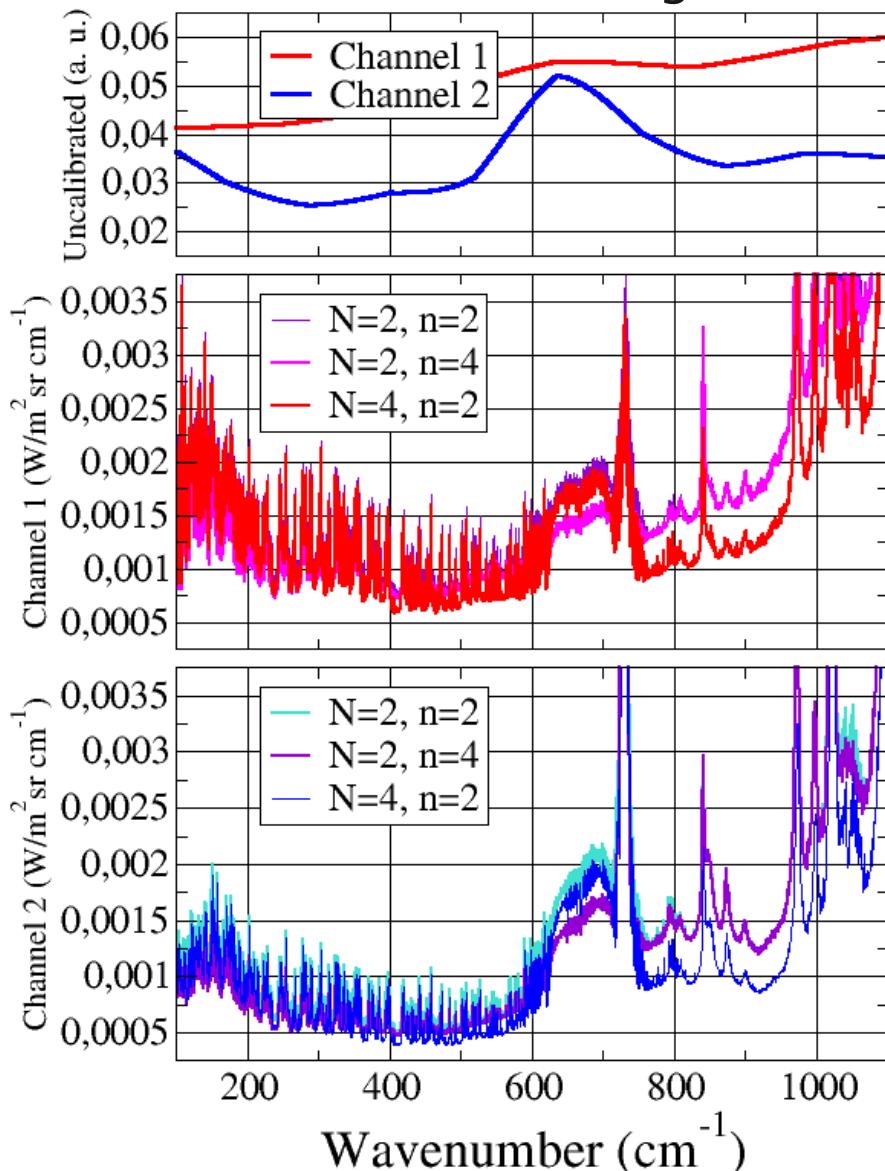
RBB  
heated



HBB/CBB  
heated

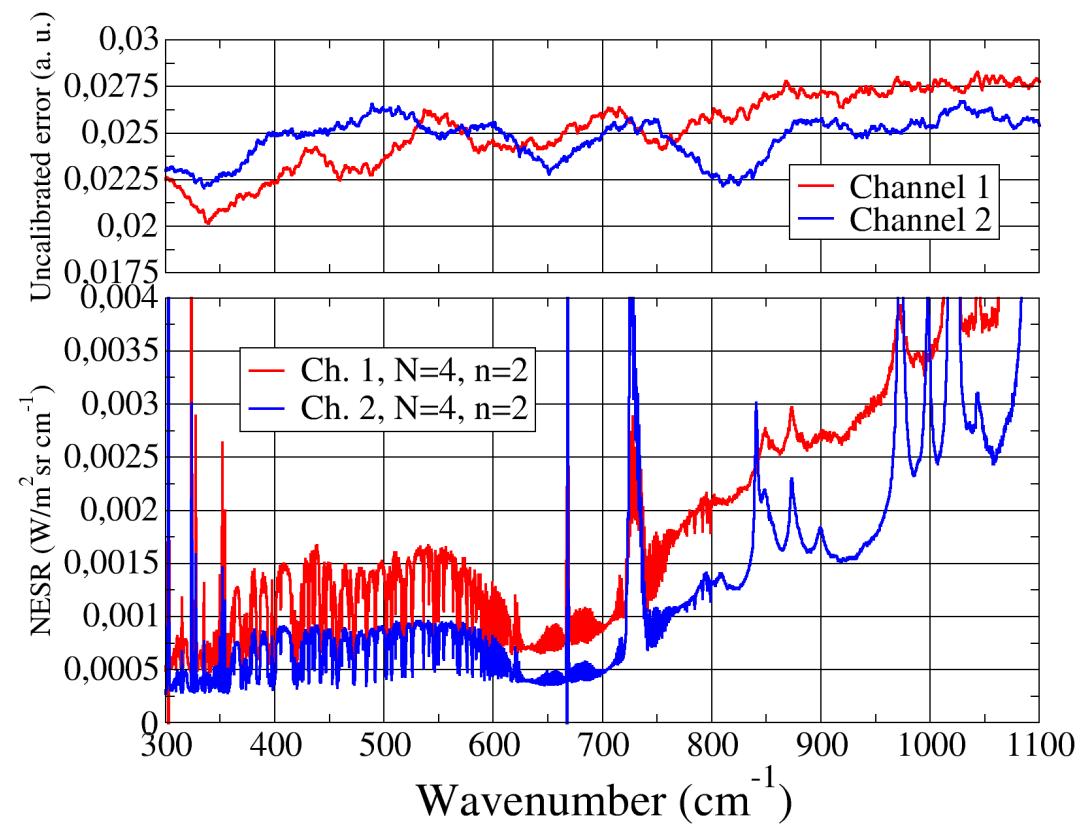


## Nadir looking



$$NESR = \sqrt{\frac{1}{N} + \frac{2}{n} \left( \frac{s}{s_h - s_c} \right)^2 \frac{\Delta S}{F1}}$$

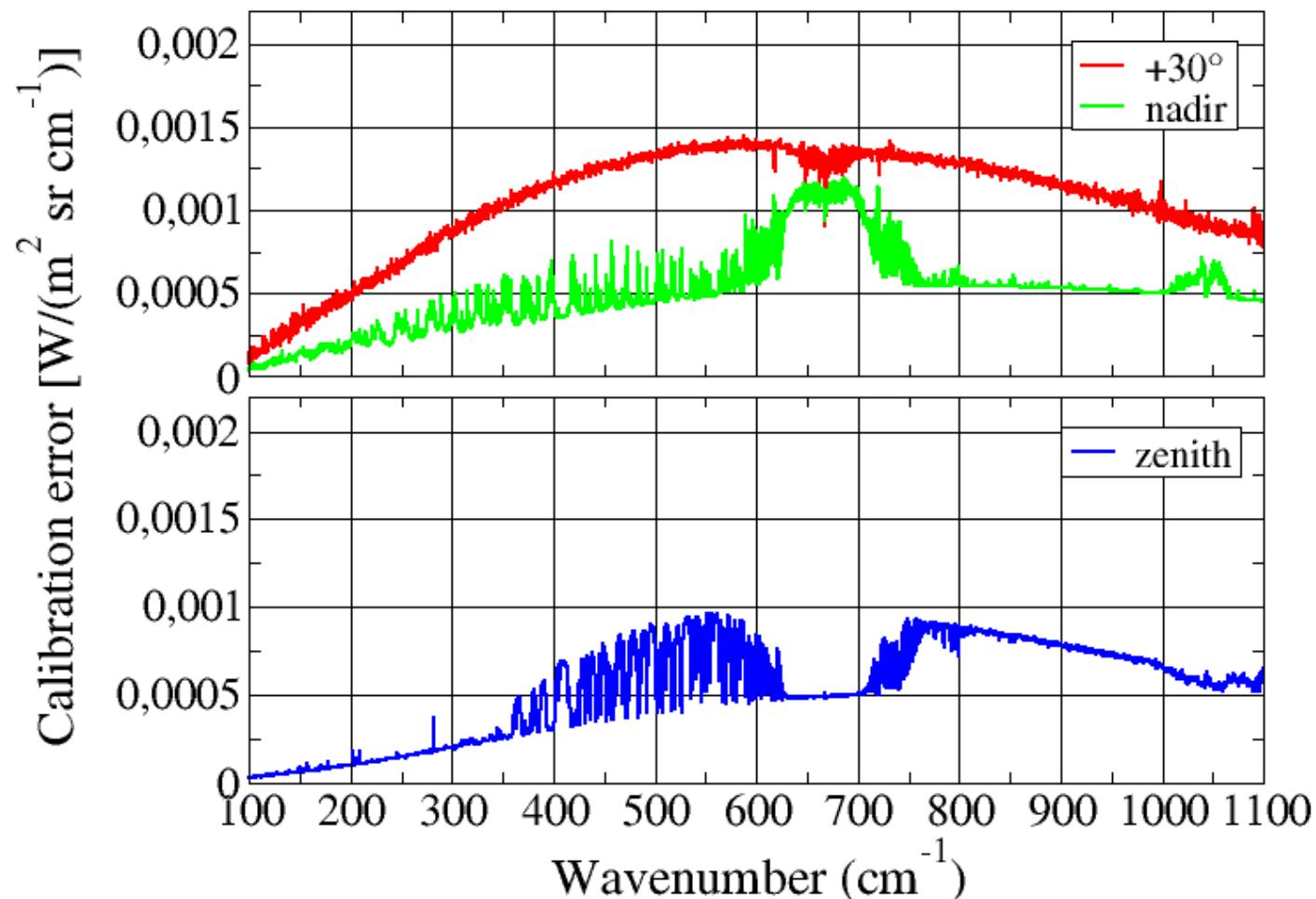
## Zenith looking



# Calibration Error

$$\Delta L = \sqrt{\Delta B_r^2 + \left( \frac{s}{s_h - s_c} \right)^2 (\Delta B_h^2 + \Delta B_c^2)}$$

Estimate with  $\Delta T_{\text{BB}} = 0.3 \text{ K}$



# The REFIR-PAD Experiment

## Balloon-borne and Ground-based Campaigns

**TABLE 1.** Data available from the measurement campaigns of REFIR-PAD

	Date	Time (UTC)	Meas. time (min)	Spectral band (cm <sup>-1</sup> )	Resol. (cm <sup>-1</sup> )
Teresina, Brazil *	30 June 2005	8:05–15:48	6.4	100–1100	0.475
Monte Morello, Italy †	6 February 2006	16:26–17:58	5.1	350–850	0.5
Monte Gomito, Italy **	13 March 2006	16:20–9:30 (+1 d)	6.1/9.9	350–1100	0.5
Testa Grigia, Italy ‡	4 March 2007	19:20–23:39	5.1	240–1400	0.5
	5 March 2007	17:54–0:43 (+1 d)			
	9 March 2007	7:25–13:53			
	11 March 2007	16:22–2:06 (+1 d)			
	12 March 2007	8:44–15:45			
	12 March 2007	17:55–23:02			
	13 March 2007	9:15–14:03			
	13 March 2007	18:21–8:04 (+1 d)			
Breuil-Cervinia, Italy §	15 March 2007	15:14–23:09	5.1	350–1400	0.5
Pagosa Springs, USA ¶	22 April 2009	17:15–19:58	5.1	350–1400	0.5
	23 April 2009	22:24–23:21			
	24 April 2009	11:30–17:58			
	25 April 2009	12:33–18:23			
	27 April 2009	12:34–18:05			
	28 April 2009	14:09–17:59			
	29 April 2009	19:04–23:29			
Cerro Toco, Chile	from 21 August 2009 - - - 37 days - - - to 24 October 2009	-	5.1	100–1500	0.5

\* OLR from balloon at 34 km, 5.078° S, 42.874° W

† DLR from ground at 610 m a.s.l., 43.844° N, 11.246° E

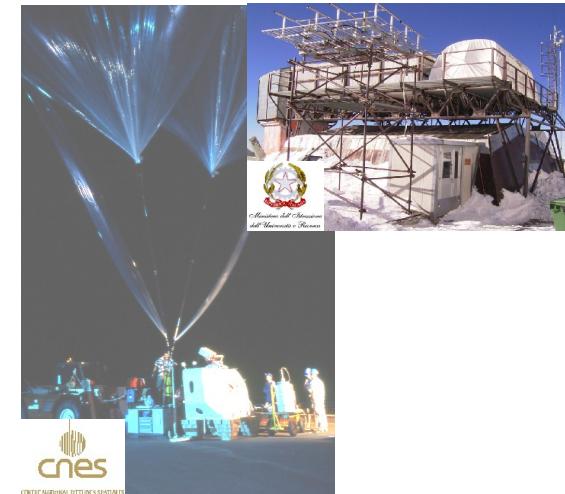
\*\* DLR from ground at 1892 m a.s.l., 44.128° N, 10.644° E

‡ DLR from ground at 3480 m a.s.l., 45.933° N, 7.7° E

§ DLR from ground at 1990 m a.s.l., 45.933° N, 7.6° E

¶ DLR from ground at 2329 m a.s.l., 37.28° N, 107.08° W

|| DLR from ground at 5340 m a.s.l., 23° S, 68° E



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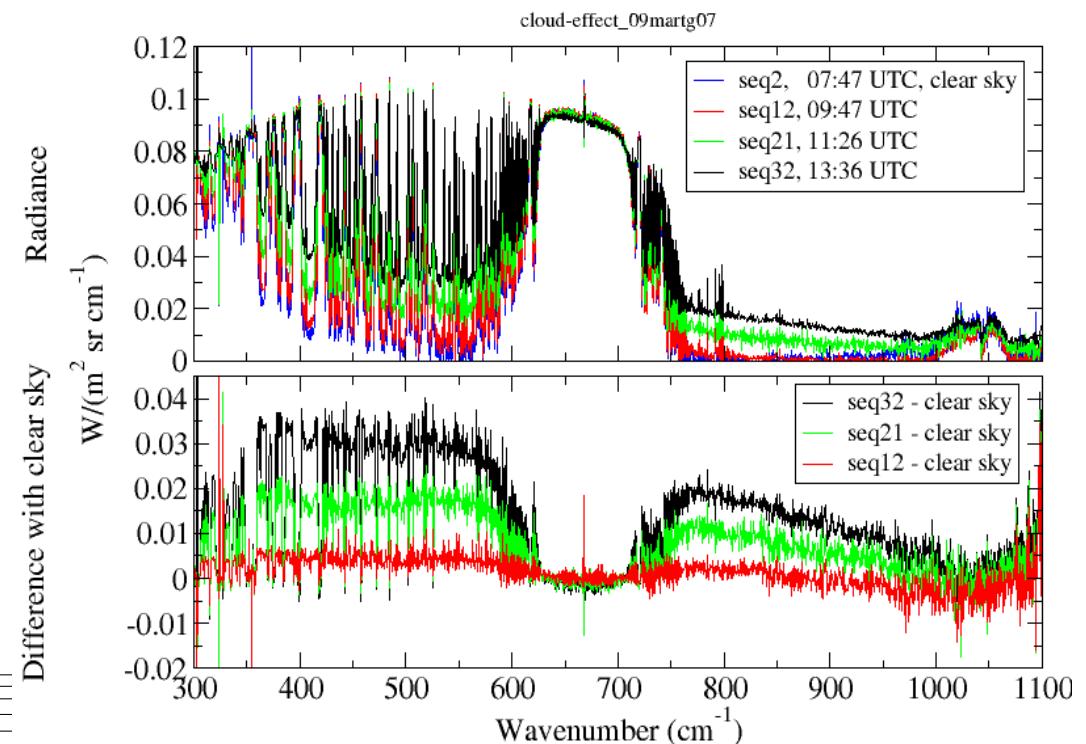
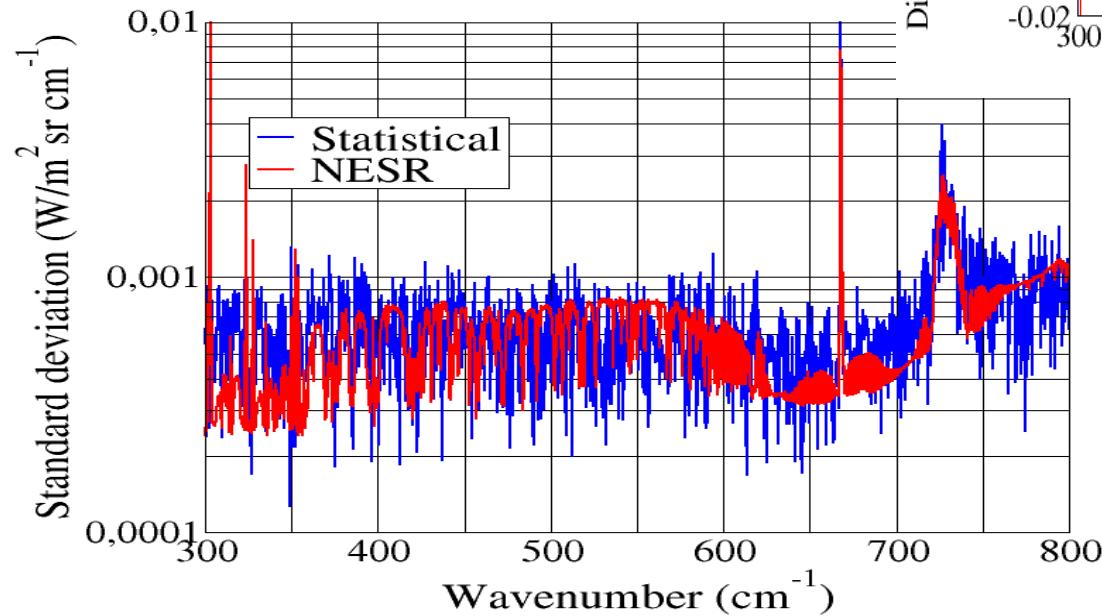
**ADGB**  
Atmospheric Dynamics Group Bologna

**ifac**

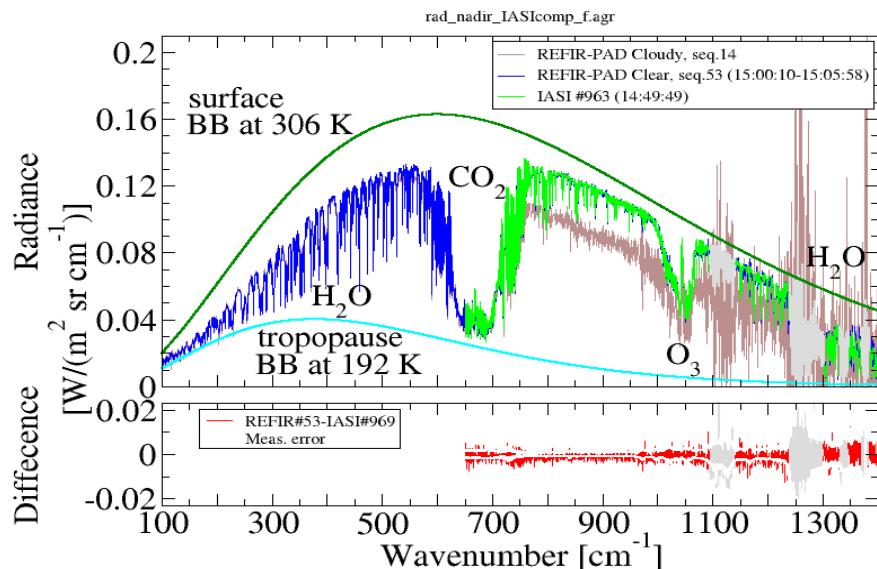
**imaa**  
Istituto di metodologie per  
l'analisi ambientale

# Atmospheric measurements and STD

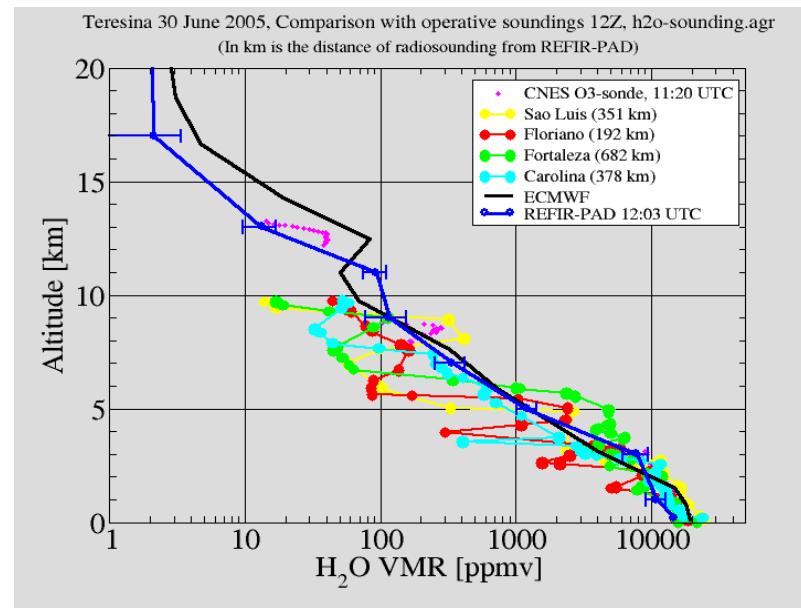
Cerro Toco 2009  
From clear sky to cirrus cover



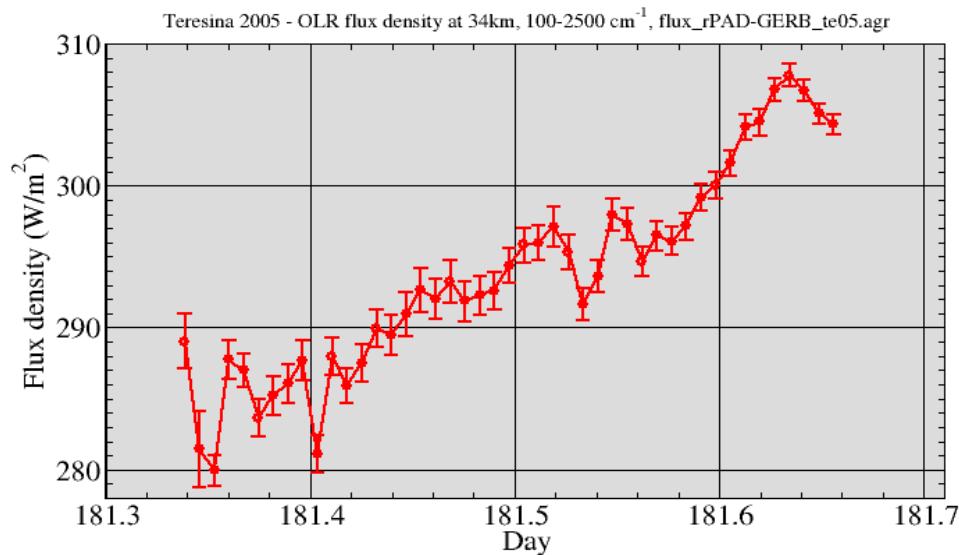
# Products



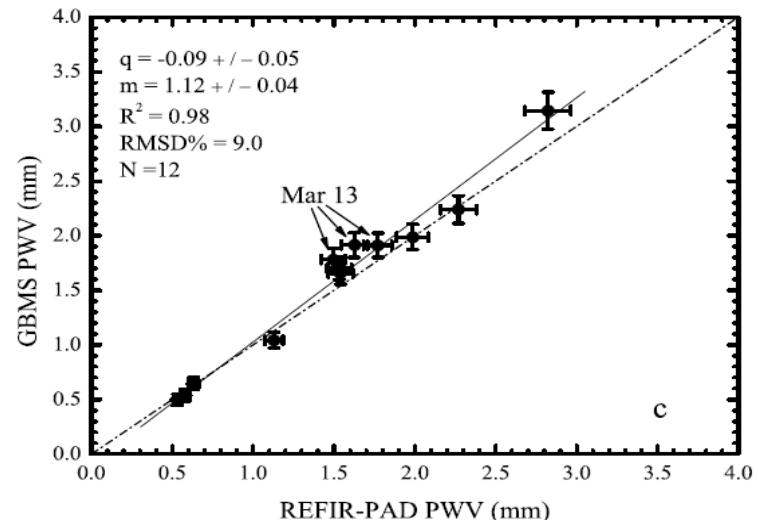
OLR accuracy = 0.3K



VMR<sub>H<sub>2</sub>O</sub> error = 22-35% from 0 to 17 km



TOA irradiance Error = 1.3 W/m<sup>2</sup>



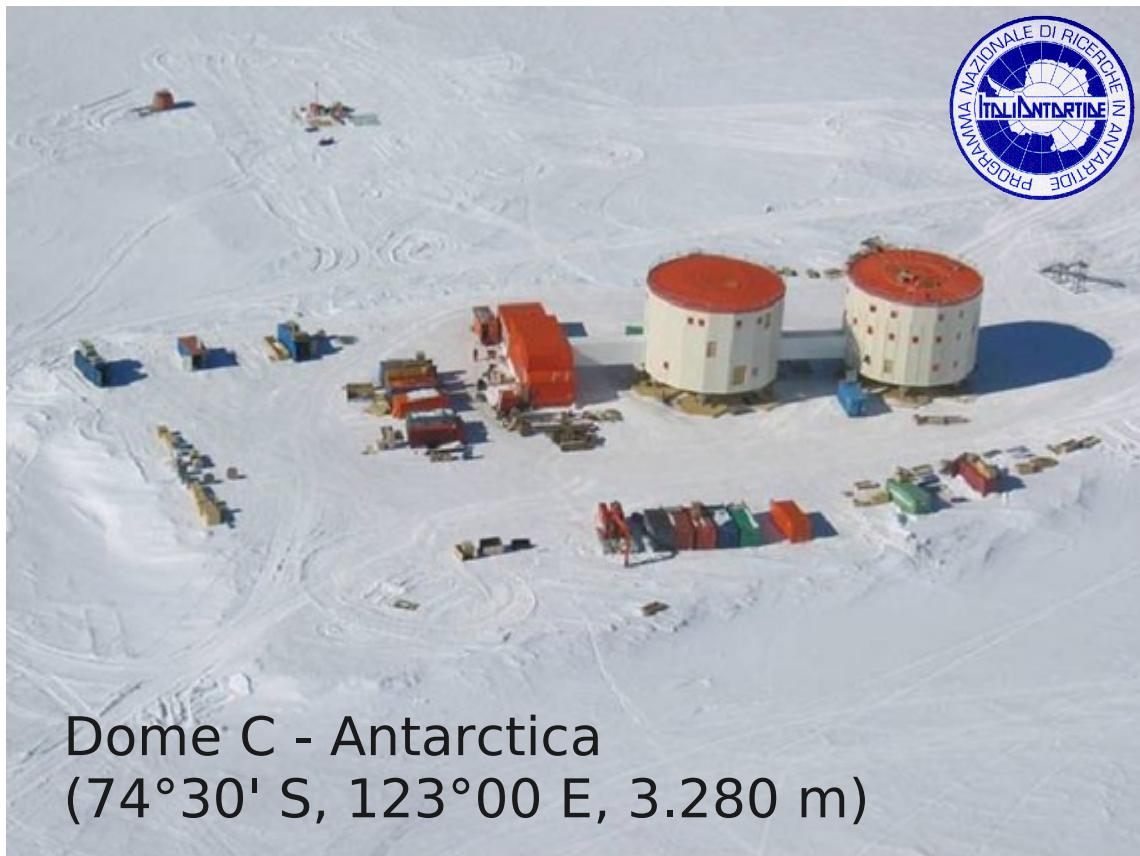
Retrieval of PWV, error = 5%

# FIR ground-based measurement opportunity

## 2011-2013

- Deployment of REFIR-PAD at the Italian-French station of Concordia on the Antarctic plateau (Dome-C) on Dec. 2011

PNRA-Programma Nazionale di Ricerche in Antartide



## PRANA project

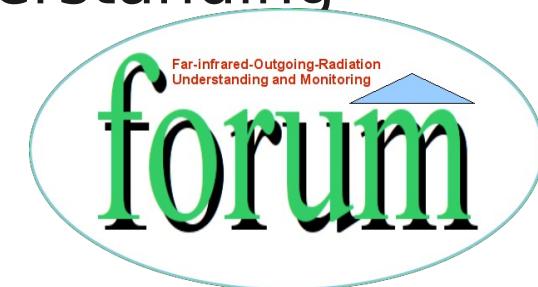
- Scientific Objective
  - Study of the radiative properties of WV and clouds in the FIR spectral region
- Available instruments
  - REFIR-PAD
  - Backscatter LIDAR
  - Radiosoundings

# FORUM and EFTWVAC

## Main mission objectives



- Study of the forcing/feedback effect on the climate system of the atmospheric water, in the form of both vapour and clouds, by measuring from space on a global scale for the first time the spectrally-resolved emission of the Earth in a broad spectral range that includes the FIR region.
- Far Infrared Outgoing Radiation Understanding Monitoring
  - Polar satellite
- Emission Fingerprints of Tropical Water Vapour and Clouds
  - ISS



**EFTWVAC**

# Space mission opportunities

## FORUM - 2010-11

ESA – Earth Explorer Opportunity Mission EE8 results

- FORUM was in a short list of 4 but it was not financed. ESAC recognises “the very high scientific interest in a radiation mission, measuring the far infra-red spectrum for the first time and examining important dependencies on cirrus cloud properties”.
- ESAC committee recommended that ESA initiate a study, to better identify the benefits of a FORUM-type mission (wavelength coverage, radiometric performance, etc.). This is under investigation.

## EFTWVAC - 2011-12

ESA – AO for ISS Experiments relevant to study of Global Climate Change

- A national support for the deployment on the ISS is also under investigation.



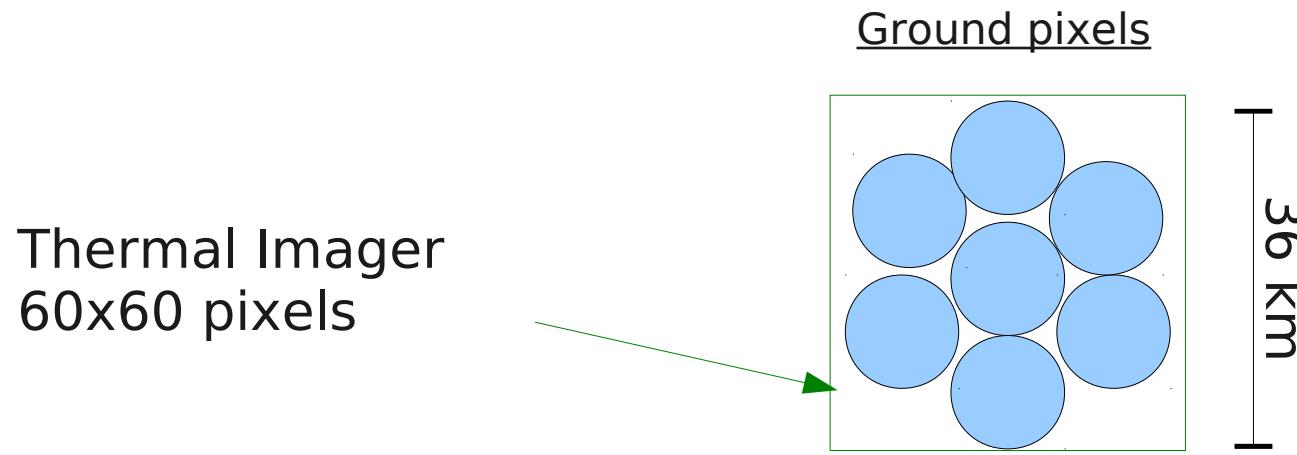
# Specific objectives

## Instrument requirements

- Spectrally resolved observation of the OLR for the attribution of the changes of total Earth irradiance to the underlying climatic parameters ( $H_2O$ ,  $CH_4$ ,  $O_3$ , etc.)
  - Spectral range =  $100-1600\text{ cm}^{-1}$
  - Radiometric accuracy =  $0.1\text{ K}$
  - Observing mode = nadir
- Determination of the atmospheric state (improved WV profiles in the upper troposphere) and assessment of its relationship with the LW spectral radiance and irradiance.
  - OPD =  $\pm 2.5\text{ cm}$  ( $0.2\text{ cm}^{-1}$  resolution)
  - NESR =  $0.2\text{ mW/m}^2\text{-sr}\cdot\text{cm}^{-1}$
  - Max resolving power, Rmax = 2500
- Improved cloud characterisation using the new information present in the FIR, and assessment of the LW contribution of clouds to the ERB.
  - Ground pixel =  $12\text{ km}$
  - Thermal imager ( $10.5-12.5\text{ }\mu\text{m}$ ) for identifying pixel contamination

# Imaging capability

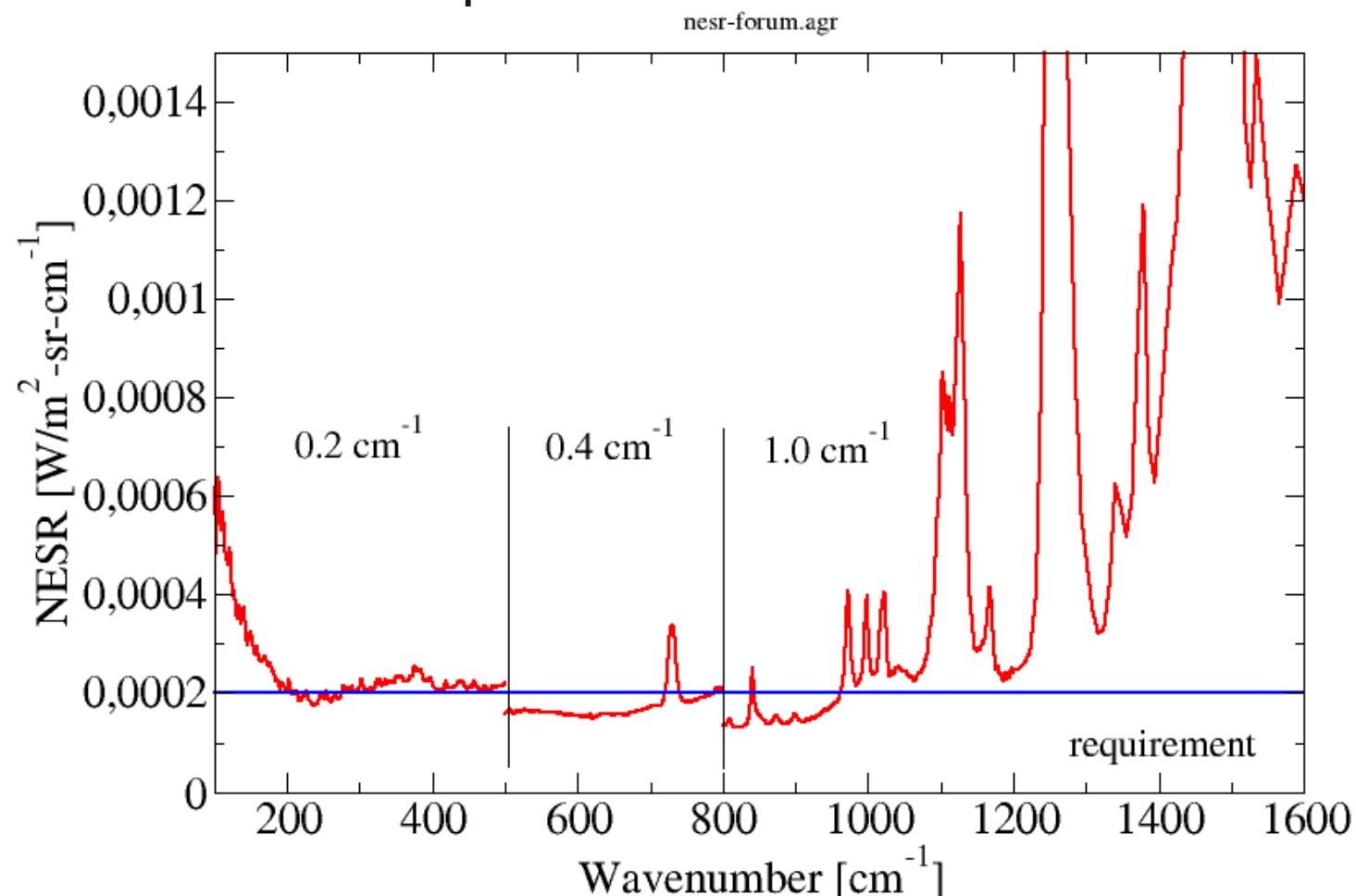
- Resolution requirements
  - $0.2 \text{ cm}^{-1}$  ( $\text{H}_2\text{O}$  retrieval) in  $100\text{-}500 \text{ cm}^{-1} \rightarrow R_{\max} = 2500$
  - $0.4 \text{ cm}^{-1}$  (T retrieval) in  $500\text{-}800 \text{ cm}^{-1} \rightarrow R_{\max} = 2000$
  - $1 \text{ cm}^{-1}$  (fingerprints) in  $800\text{-}1600 \text{ cm}^{-1} \rightarrow R_{\max} = 1600$
- Aperture limit on  $R_{\max}$  for a FOV = 7 pixels (optical throughput of single pixel =  $0.01 \text{ cm}^2 \text{ sr}$ )
  - Limit on  $R_{\max}$ -central = 17000
  - Limit on  $R_{\max}$ -off axis = 2100



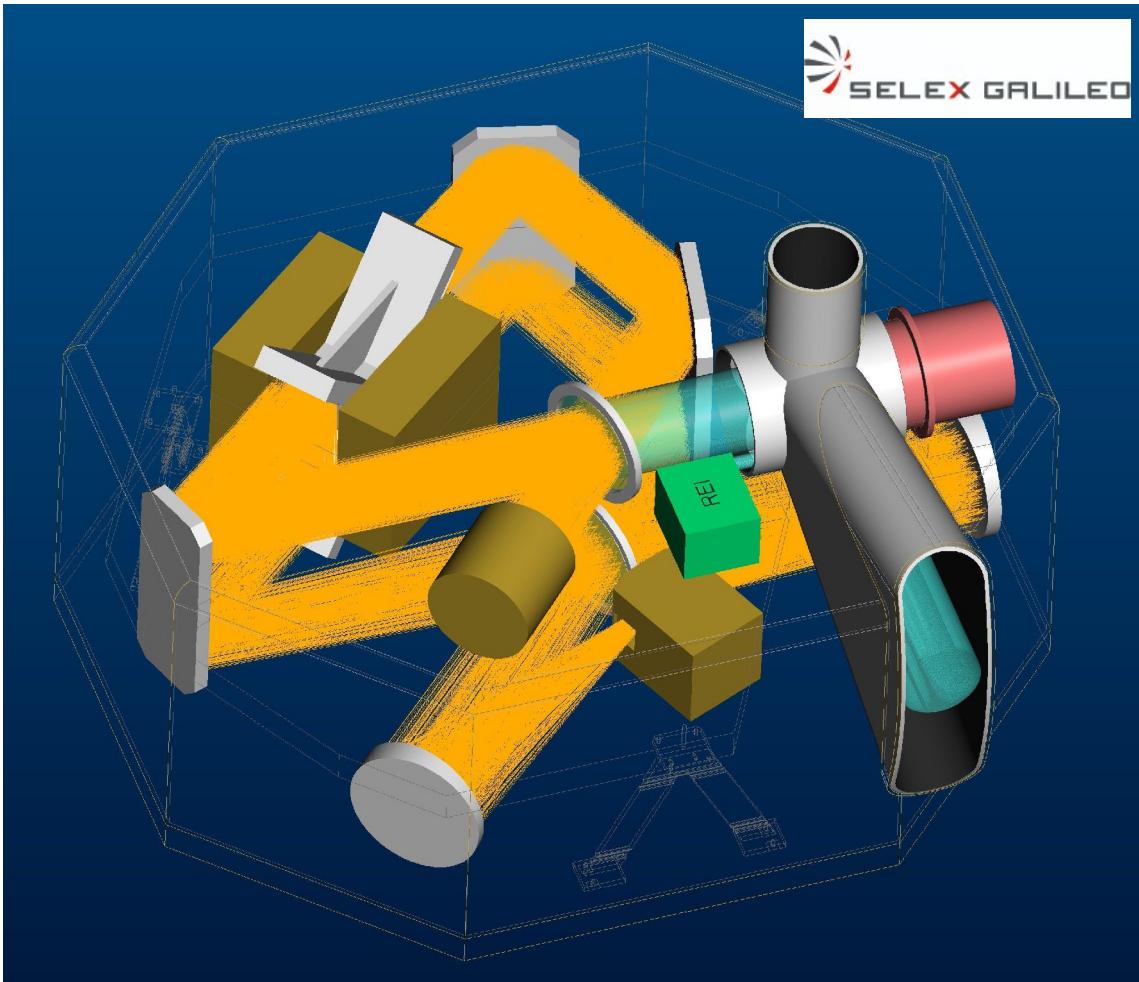
# Performances

## Expected radiometric precision

- NESR requirement (blue curve) compared to the estimation performed for the FIR FTS of FORUM/EFTWVAC based on REFIR-PAD performances



# FTS optomechanical design



1 m

- FTS specifications
  - 14 uncooled pyroelectrics ( $D^* \approx 10^9 \text{ cm}\sqrt{\text{Hz/W}}$ )
  - Sampling rate = 1.25 kHz
  - Acquisition time = 32 s
  - Weight = 70 kg
  - Power = 40 W
  - $T_{\text{BB}}$  abs. cal. = 50 mK
- EI = Embedded Imager

# Acknowledgements and References

## • Acknowledgements

- Financial support for ESA proposal provided by the project POR-CREO-FESR-2007-2013, CTOTUS of the Italian Regione Toscana.



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